

Not afraid to ride the wind: A Montgolfière for Titan

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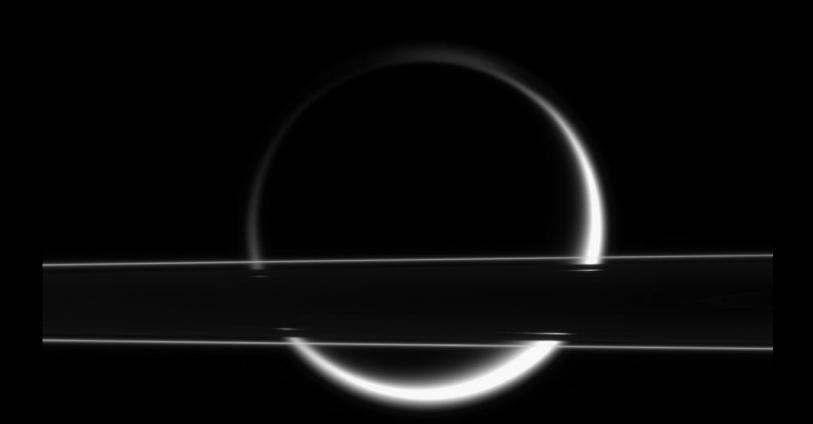
Ralph Lorenz

and the TiPEx Team

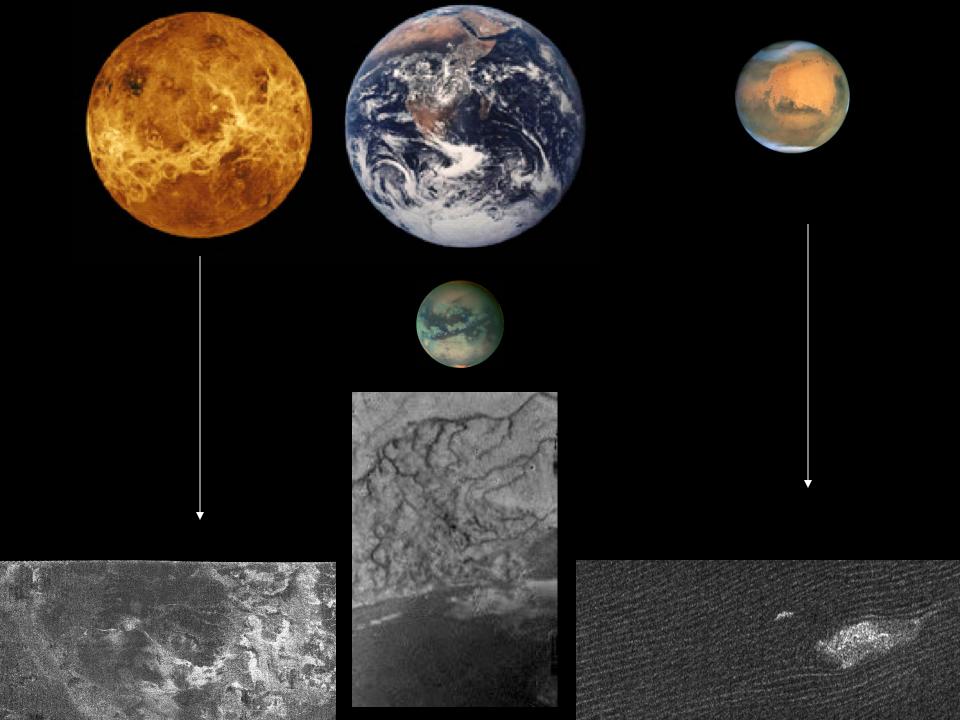


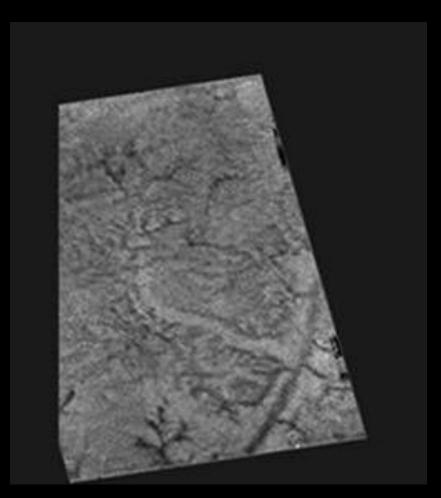


Titan is big....

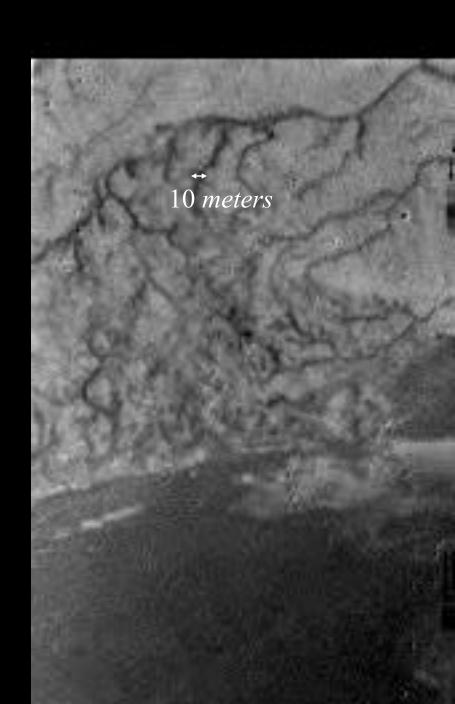


Titan's atmosphere is dense, nitrogen-methane, cold...



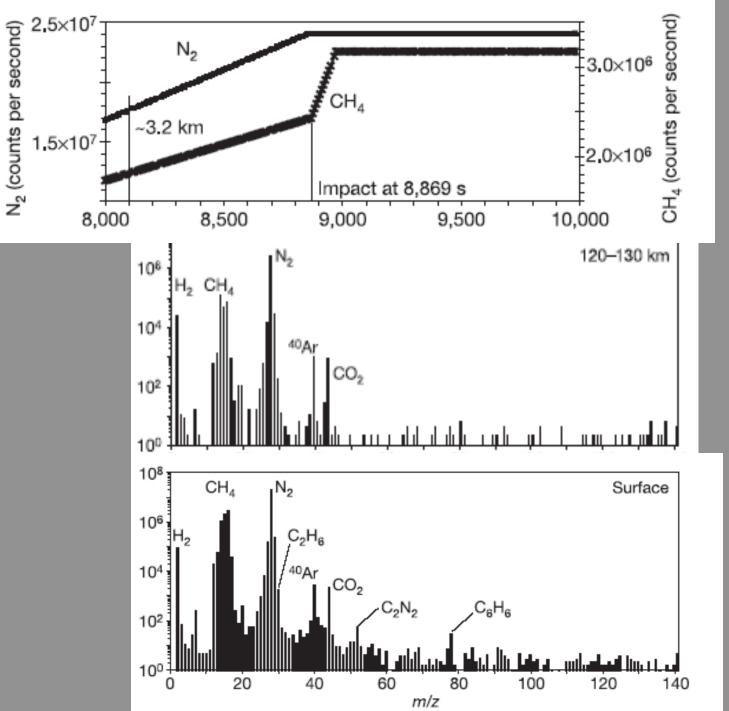


Rains and springs of methane carve channels and widen fractures in the water ice crust...

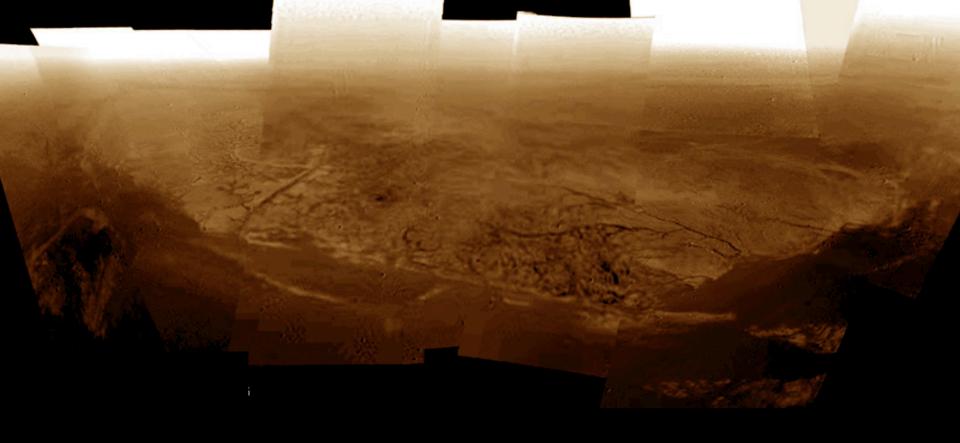




...and dump pebby debris on surrounding plains







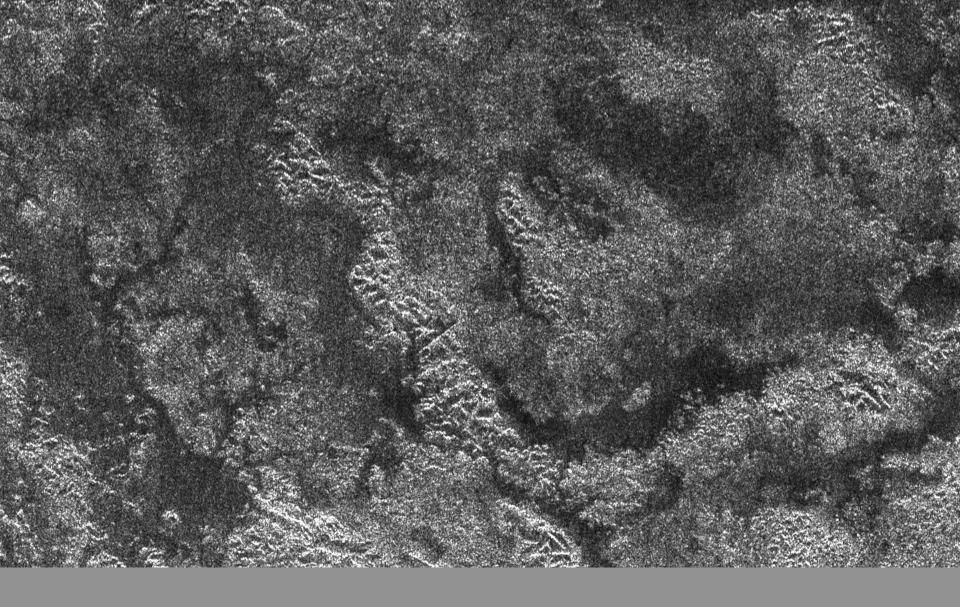
Titan has vistas that evoke—more than any other world beyond Earth—fantastical *lands* of the human imagination...



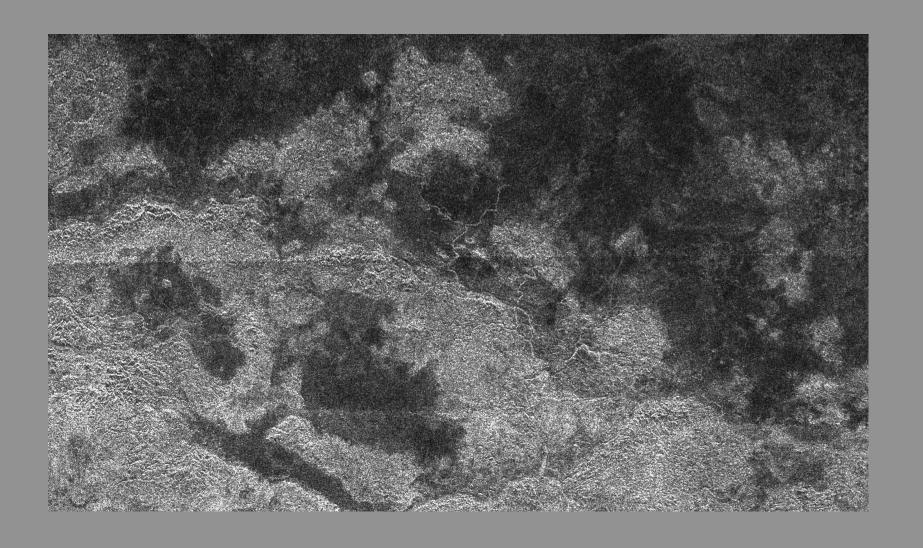
"In Xanadu did Kubla Khan A stately pleasure-dome decree:



...where Alph, the sacred river, ran...



...through caverns measureless to man...



...down to a sunless sea."

Samuel Taylor Coleridge, Xanadu

"Titan is a complex world, which appears to be influenced by tectonic, fluvial and atmospheric processes, in many ways similar to Earth," *Porco et al., Science, 2005.*

"As an environment, Titan certainly meets all of the stringent criteria outlined above for life."

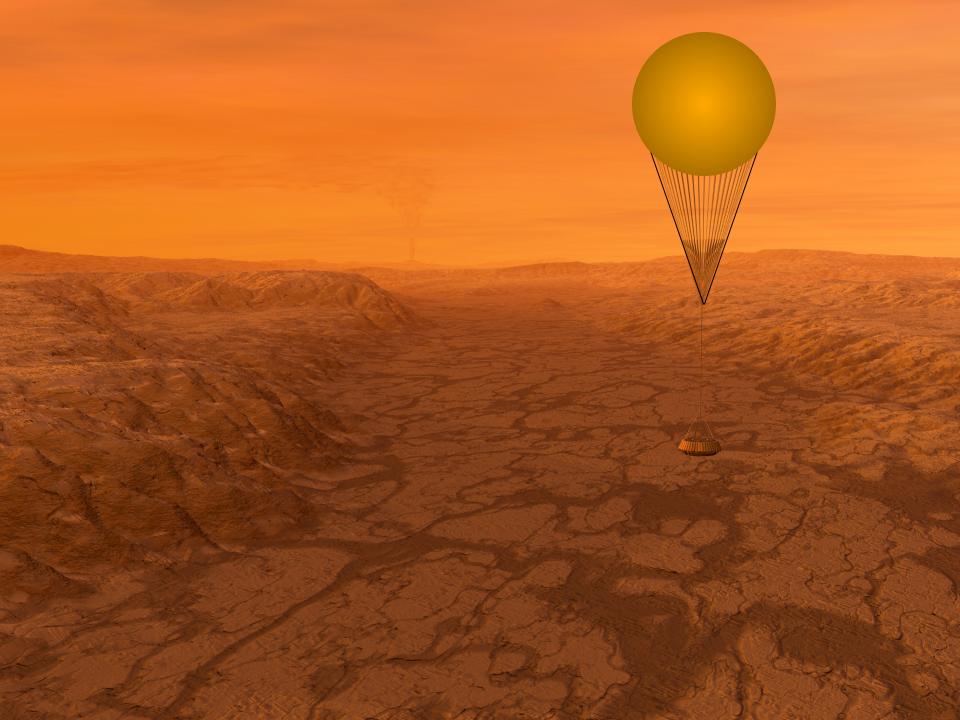
Steven Benner et al. Current Opinion in Chemical Biology, 2004

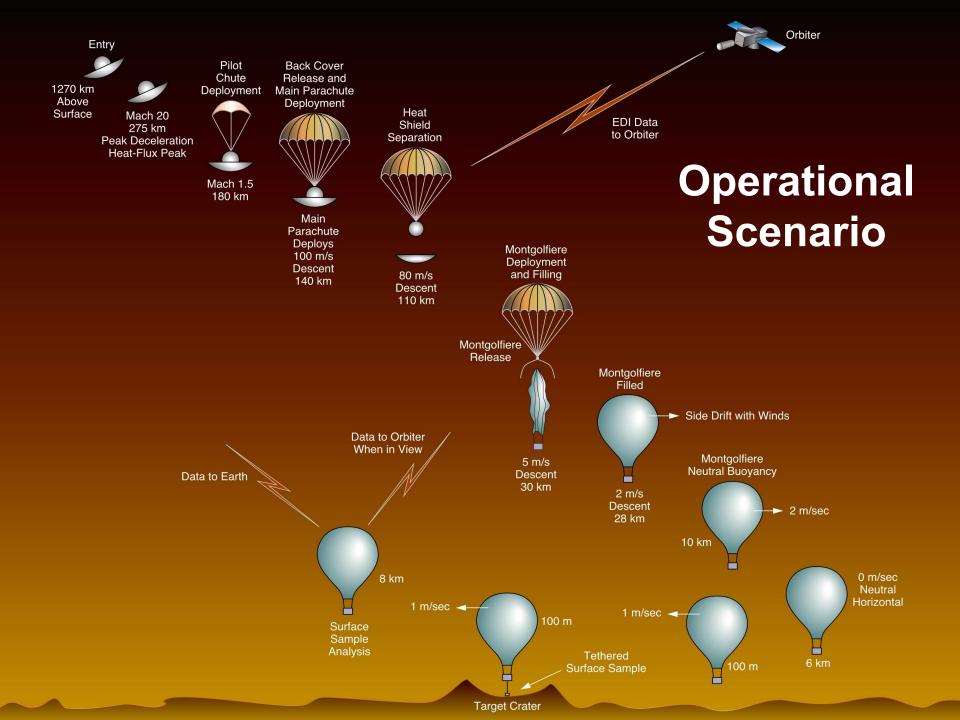
"Titan is the Peter Pan of the solar system." Tobias Owen

"It cannot now be predicted whether Europa or Titan will ultimately prove to be the most promising satellite for long-term exploration." NRC Decadal Survey of Solar System Exploration, 2003

Titan wants us!

- Low heat load entry; large scale height atmosphere
- Low gravity/high density/cold/high molecular weight environment—optimal for flying
- Benign environment: predictable or learnable winds with low velocity at cruising altitudes (10 km), no radiation, low topography (< 0.5 km; km's in places).
- Radar+DISR is telling us that there are endless evocative vistas to image at 1-10 meter resolution...





Science Goals and Measurement Objectives of the Titan Prebiotic Explorer (Tipex)

- Determine the origin and evolution of Titan's atmosphere
 - Extent of internal differentiation of Titan's interior
 - Surface evidence for primordial atmospheric composition
 - Differentiation in isotopic ratios between primordial surface constituents and atmospheric constituents
- Determine the cycling of carbon through the interior, surface and atmosphere on all timescales and in all phases
 - Titan methane meteorological cycle
 - Chemical modification of carbon speciation over geological time
 - Sources of methane
- Characterize the prebiotic chemistry that has occurred in the surface and immediate subsurface
 - Extent of chemical evolution of surface organic deposits in geologically active and inactive zones
 - Elemental and mineralogical composition of surface inorganics in contact with organics

Science Traceability to Instruments

		Orbiter					Aerial Vehicle						Surface Sample		
		Imager/altimeter	Spectrometer	Radio Science	Subsurface Near-IR Imaging Spectrometer		Aeriai veriic		licie	Temperature		det	det Age Date IIIst Acchieve		
		(Rima) λ = TBD (between 0.6-	(NIRIS-O)	(RaSci)	(SubS)	(NIRIS-A)	Spectrometer (TDLS)	(SoA)	(IC)	Sensor (PTS) 1500 - 1600 mBar with	(GCMS)	(GCMS)	(AD)	(EA)	
Measurement Objectives		λ = TBD (between 0.6- 6 cm), TBD m antenna diameter;	λ = 1-5 micron, $λ/Δλ =200; spatial resolution100 meters$	Frequency of TBD, USO of TBD perf on Sonde	Long-wavelength (~10 meters) pulse-sounder system, TBD peak	λ =1-5 micron, λ/Δλ= 200; spatial resolution 100 meters, FOV TBD	TBD wavelength range see Chris Webster	for time varying winds	Upward and downward looking (+- 60 degrees FOR) stereo imaging		Neutral & ion species, mass range = 1-600 amus with 5 ppm	Neutral & ion species, mass range = 1-600 amus with 5 ppm	Geiger counter	TBD instrument parameters and values	
		simultaneous spatial imaging and altimetry			output power, TBD duty cycle	or λ = 0.6-1.5 microns broadband; spatial resolution 1 meter;		and T. P,T to small errors (0.1 K in T, ?? In	cameras; broadband		resolution.	resolution.			
						Imaging mode 10 cm resolution FOV TBD		r).	cm best resolution						
Science Goals Data Products Products															
		10% of total surface area mapped at 100 m	Global spectral hypercubes in 6	Orbit tracking with pulse altimeter	Echo Profiles with depths up to several	Spectral hypercubes in 6 selected wavelength	Isotope abundances	Wind velocities with TBD accuracy and	Imaging data set of 20 km x 20 km survey	Pressure and temperature history	Gas-phase abundances of neutral	Abundances of neutral and ion species with	Relative abundance of Carbon 14 surface	identify elemental composition of surface	
		spatial resolution; altimetric overlays to + 10 m	selected wavelength bands of interest x 5 channels per band;	geodetic overlays yields 10 meters precision in Orbiter	hundred meters with 1m depth resolution.	bands of interest x 5 channels per band; 400 Gbit data volume		precision and cloud maps; P,T over TBD area with temporal	boxes; data volume 1 Gigabit per box (3 x compression); 1 box		and ion species with masses between 1 and 600 amus	masses between 1 and	sample	sample surface sample	
		10 111	400 Gbit data volume with 10x image	position over time. Sonde location to		with 10x image compression; Imaging		resolution TBD	per week.		600 amus				
			compression; 100 meters best resolution	within 100 meters		data set of 20 km x 20 km survey boxes; data volume 1 Gigabit per									
						box (3 x compression); 1 box per week.									
		_								^-l					
Science Goals Measurement objectives	Required measurements ("levels 3 & 4" data products) Time varying gravitational moments; k2 to an accuracy of 0.001; J2 to an	-													
Extent of internal differentiation of Titan interior Surface evidence for primordial tetermine the origin and evolution of tarns atmosphere	accuracy (absolute) of 10-9. [TBD because the thick atmosphere complicates matters—check with Julie Castillo at JPL as to what is possible.]			х											
	Global distribution of the abundance of ammonia and other "cryovolcanic" compounds on surface with spatial resolution of 10's to 100's of meters. Sensitivity of 0.1%		x			х									
Differentiation in isotopic ratios between primordiral surface	12C/13C and 14N/15N in atmospheric deposits of methane, ethane, HCN, acetylene, ethylene, carbon dioxide, ammonia, nitrogen to an accuracy of 1%						x					x			
constituents and atmospheric constituents	12C/13C and 14N/15N in surface methane, hydrocarbons and nitriles, to an accuracy of 1%.						х					х			
Titan methane meteorological cycle etermine the cycling of carbon trough the interior, surface and troughpiere on all timescales and in it phases Chemical modification of carbon speciation over geological time	Map of fluvial features to meter and decimeter spatial resolution and context with aeolian features over an area of a few hundres square km					х			х						
	TBD cloud particle size distribution, shape and density Cloud appearances and disappearences with spatial resolution of 100s of						X								
	meters Methane mixing ratio profile at TBD vertical resolution to better than 0.5%	4	х						Х						
	Winds measured at TBD frequency of >0.5 Hz							Х							
	Temperature profiles to 0.1 degree Subsurface structure to depth of several 100 meters with 1m depth resolution	1			х					Х					
	over prescribed survey boxes of 100s of square km. Molecular composition of the surface at 100s of square meter resolution;	-			^										
	Molecular composition of the surface at 100s of square meter resolution ; selected sites at 10 meter resolution		Х			х									
	Global radar of the horizontal structure of the surface at 100s of square meters spatial resolution coregistered with composition measurements	х													
	Global topography at 100s of square meters spatial resolution and 10 m vertical resolution coregistered with composition and structure measurements	x													
	Images of layering in fluvial deposits at centimeter spatial resolution to a depth of 1-5 m								х						
	Abundances of carbon bearing molecules in gas (100 amu +- 0.2 amu mass accuracy and a resolution of 500 m/delta-m with a sensitivity abundance of TBD PPB to accuracy of 10%) and condensed phases (600 amu 0.2 amu										x				
	mass accuracy and a resolution of 5 PPM with a sensitivity abundance of TBD PPB to an accuracy of)														
Extent of chemical evolution of surface	Fluorescence images at spatial resolution of TBD square meters may not be achievable at reasonable altitudes due to Titan atmosphere							х							
organic deposits in geologically active and inactive zones haracterize the prebiotic chemistry	Composition of surface organics 600 amu +- 0.2 amu mass accuracy and a mass resolution of 5 PPM with a sensitivity abundance of 1 PPB to an accuracy of 10%											х			
naracterize the preblotic chemistry hat has occurred in the surface and nmediate subsurface	Ages of surface organic deposits with TBD parameters Morphology of surface sample with TBD mm resolution registered within 10												Х		
Elemental and mineralogical composition of surface inorganics in	square centimeter image of the sample site to 1 mm resolution Surface sample site characterized over 10 square centimeters at 0.1 mm	_								х					
contact with organics	spatial resolutions Abundances of the elements (Li, Na, K, Mq, Ca, Fe, Ni, S, P, F, Cl, Si, B, Al)	-							х					x	
	to > 1ppm to an accuracy of 10%	_		<u> </u>										^	

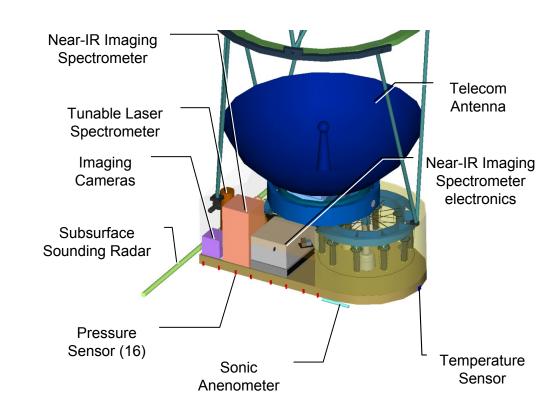
Instruments

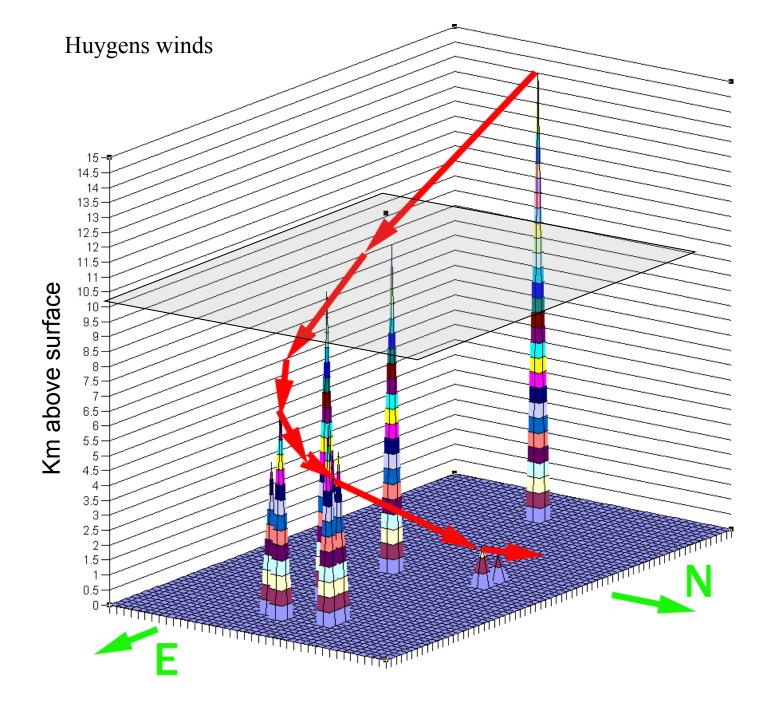
Gondola

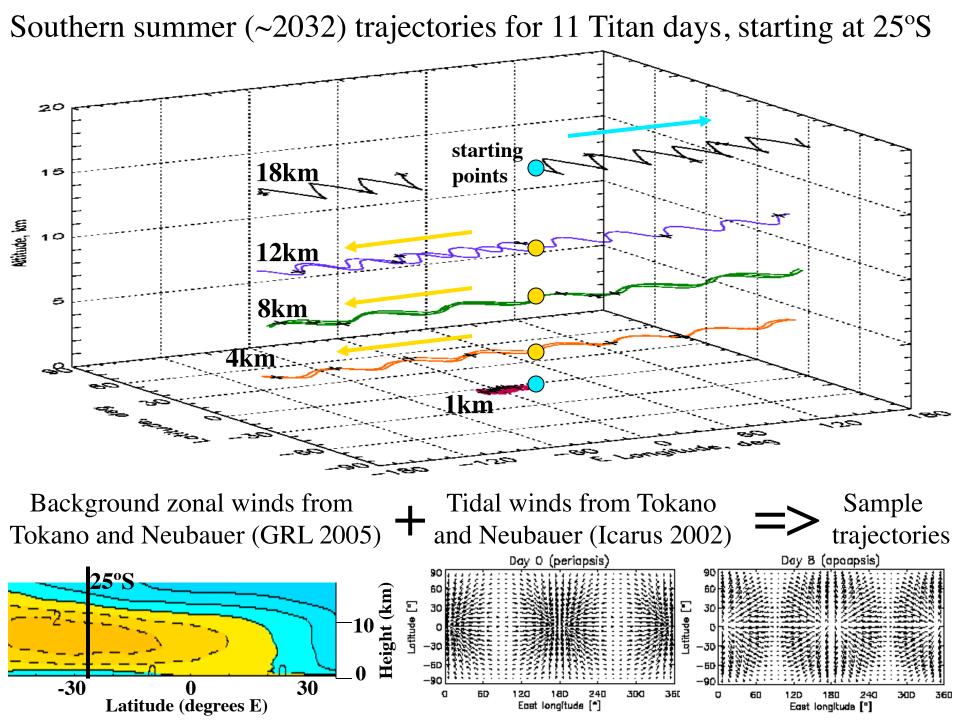
- Subsurface sounder
- Near-IR spectrometer
- Tunable laser
- Sonic anemometer
- Imaging camera(s)
- P-T sensor
- Gas Chromatograph and Mass Spectrometer
- Surface Sample Analysis
 - Sample acquisition mechanism
 - GCMS with chiral support
 - Age dating (¹⁴C detection)
 - Surface hardness
 - Sample context imager
 - Sample microscope imager
 - Elemental Analysis

Orbiter

- Radar imager/altimeter
- Near-IR imaging spectrometer
- Radio science

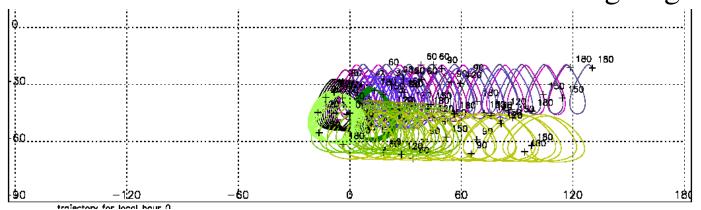




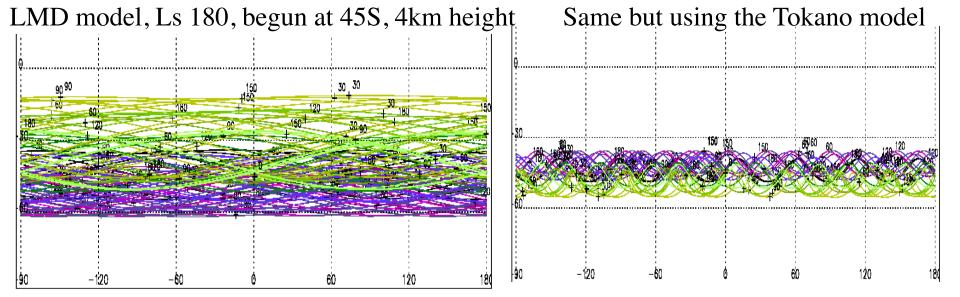


The impact of starting at different local times

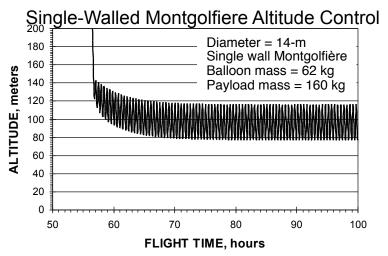
Each color is a different local start time (12 total spanning an entire Titan day) - notice how both the amount of wander from the initial location *and/or* the latitudes covered can change significantly



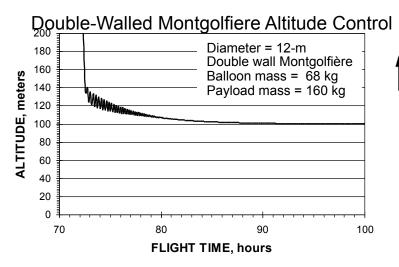
Tokano model, Ls 270, begun at 45S, 18km height

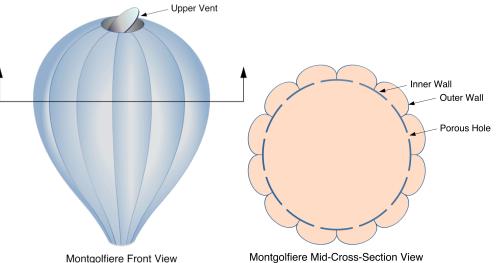


Titan Montgolfière Thermal Analyses



- Both the single-walled and double-walled RPS Montgolfière can function well on Titan.
- Double-walled Montgolfière can hold tighter altitude control and can be smaller (12-m vs. 14-m diameter)
- Double-walled Montgolfière can safely ascend and descend at speeds of 1 km/hr. This corresponds to a 15 degree hazard avoidance angle with surface winds ~1 m/sec





Data Return Capability

Orbiter

- Ka-band DTE provides 200 kb/s (power and pointing limited)
- Occultations 0-35% depending on orbit geometry
- 1-10 Gb/day
- Assumes worst case geometry and continuous DSN coverage
 - Edge-on geometry is worst case with 35% occultation

Aerobot

- X-band with 1m HGA
- 2 kb/s DTE
 - Useful for commanding, navigation Doppler, health and status, low rate science for sequencing and dead-reckoning
- 1 Mb/s relay to orbiter
- 5 Gb/day average
- Assumes worst case geometry
 - Low latitude location is worst case with relay contacts 3 days in 8

Technology developments

- Low-T materials
- Adaptive navigation autonomy
- Aerocapture (orbiter)
- Miniaturization for instrument/sampling

International partnering ideas

- Orbiter and balloon US; Gondola ESA
- Orbiter ESA; Balloon/gondola US
- Balloon/gondola US; Enceladus package ESA

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To explore strange new worlds....

